


# COLLECTIVE EFFECTS

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USPAS Fundamentals, June 4-15, 2018

E. Prebys, Accelerator Fundamentals: Collective Effects

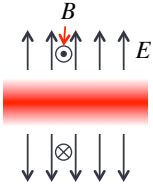
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## Space Charge

So far, we have not considered the effect that particles in a bunch might have on each other, or on particles in another bunch.

Consider the effect of space charge on the transverse distribution of the beam.



$$\rho(r) = \frac{Ne}{2\pi\sigma^2} e^{-r^2/2\sigma^2}$$

radial charge density

If we look at the field at a radius  $r$ , we have

$$\oint \vec{E} \cdot d\vec{A} = E(2\pi rL) = \frac{Q_{encl}}{\epsilon_0} = \frac{Ne}{\sigma^2} \int_0^r r e^{-r^2/2\sigma^2} dr$$

$$= Ne \left(1 - e^{-r^2/2\sigma^2}\right)$$

$$\rightarrow \vec{E} = \frac{Ne}{2\pi\epsilon_0 rL} \left(1 - e^{-r^2/2\sigma^2}\right) \hat{r}$$

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Similarly, Ampere's Law gives

$$\oint \vec{B} \cdot d\vec{l} = 2\pi r B = \mu_0 I_{\text{enclosed}} = \mu_0 \frac{Nev}{\sigma^2 L} \int_0^r r e^{-r^2/2\sigma^2} dr$$

$$\rightarrow \vec{B} = \mu_0 \frac{Nev}{2\pi r L} (1 - e^{-r^2/2\sigma^2}) \hat{\theta}$$

$$\rightarrow \vec{F} = e(\vec{E} + \vec{v} \times \vec{B})$$

$$= \frac{Ne^2}{2\pi L} (1 - e^{-r^2/2\sigma^2}) \left( \frac{1}{\epsilon_0} \hat{r} + v^2 \mu_0 (\hat{s} \times \hat{\theta}) \right)$$

$$= \hat{r} \frac{Ne^2}{2\pi r L \epsilon_0} (1 - e^{-r^2/2\sigma^2}) (1 - \beta^2)$$

$$= \hat{r} \frac{ne^2}{2\pi r \epsilon_0 \gamma^2} (1 - e^{-r^2/2\sigma^2}); \quad n \equiv \frac{N}{L} = \frac{dN}{ds} \quad \text{Linear charge density}$$

$= -\hat{r}$   
 $= \frac{1}{\epsilon_0} (\epsilon_0 \mu_0) = \frac{1}{\epsilon_0 c^2}$

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We can break this into components in x and y

$$F_x = |F| \frac{x}{r}$$

$$= \frac{ne^2}{2\pi \epsilon_0 \gamma^2} \frac{x}{r^2} (1 - e^{-r^2/2\sigma^2})$$

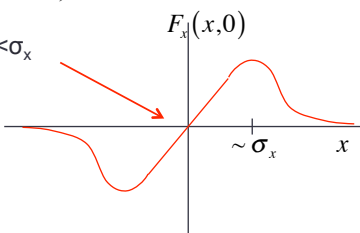
$$= \frac{ne^2}{2\pi r \epsilon_0 \gamma^2} \frac{x}{(x^2 + y^2)} \left( 1 - e^{-\frac{(x^2 + y^2)}{2\sigma^2}} \right)$$

$$F_y = \frac{ne^2}{2\pi r \epsilon_0 \gamma^2} \frac{y}{(x^2 + y^2)} \left( 1 - e^{-\frac{(x^2 + y^2)}{2\sigma^2}} \right)$$

Non-linear and coupled  $\rightarrow$  ouch! but for  $x \ll \sigma_x$

$$\left( 1 - e^{-\frac{(x^2 + y^2)}{2\sigma^2}} \right) \approx \frac{(x^2 + y^2)}{2\sigma^2}$$

$$\rightarrow F_x \approx \frac{ne^2}{4\pi \sigma^2 \epsilon_0 \gamma^2} x \quad \sim \text{linear and decoupled}$$

$$F_y \approx \frac{ne^2}{4\pi \sigma^2 \epsilon_0 \gamma^2} y$$


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$$F_x = \frac{dp_x}{dt}$$

$$\rightarrow \Delta x' = \frac{\Delta p_x}{p} = \frac{1}{p} \int F_x dt = \frac{1}{p} \int F_x \frac{dt}{ds} ds \Rightarrow$$

$$= \frac{1}{vp} \int F_x ds$$

$$x'' = \frac{F_x}{vp} = \frac{F_x}{\beta^2 \gamma m c^2}$$

$$\approx \frac{e^2}{4\pi \sigma^2 \epsilon_0 \beta^2 \gamma^3} n x$$

$$= \frac{r_0}{\beta^2 \gamma^3 \sigma^2} n x; \quad r_0 \equiv \frac{e^2}{4\pi \epsilon_0 m_0 c^2} = 1.53 \times 10^{-18} \text{ m for protons}$$

This looks like a distributed defocusing quad of strength  $d\left(\frac{1}{f}\right)/ds \equiv k = -\frac{nr_0}{\beta^2 \gamma^3 \sigma^2}$

so the total tuneshift is  $\Delta \nu_x = \frac{1}{4\pi} \oint k \beta_x(s) ds$

(“Off-momentum particles” lecture)

$$= -\frac{r_0}{4\pi \beta^2 \gamma^3} \oint n \frac{\beta_x(s)}{\sigma_x^2} ds = -\frac{r_0}{4\pi \beta^2 \gamma^3 \epsilon_x} \oint n ds$$

$$= -\frac{r_0}{4\pi \beta^2 \gamma^3} \frac{NB}{\epsilon_x}; \quad B \equiv \frac{n_{peak}}{\langle n \rangle} \leftarrow \text{“Bunching factor”}$$

$$= -\frac{NB r_0}{4\pi \beta \gamma^2 L (\beta \gamma \epsilon_x)} \left( \epsilon_{x,N} \right)$$

$$= -\frac{NB r_0}{4\pi \beta \gamma^2 \epsilon_{x,N}} \quad \text{Maximum tuneshift for particles near core of beam}$$

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## Example: Fermilab Booster@Injection

$K = 400 \text{ MeV}$   
 $N = 5 \times 10^{12}$   
 $\epsilon_N = 2 \pi \text{ mm-mr}$   
 $B = 1$  (unbunched beam)

$$\Delta \nu = -\frac{N r_0}{4\pi \beta \gamma^2 \epsilon_N} = -.247$$

This is pretty large, but because this is a rapid cycling machine, it is less sensitive to resonances

Because this affects individual particles, it's referred to as an “incoherent tune shift”, which results in a tune spread. There is also a “coherent tune shift”, caused by image charges in the walls of the beam pipe and/or magnets, which affects the entire bunch more or less equally.

This is an important effect, but beyond the scope of this lecture.

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## Beam-beam Interaction

If two *oppositely* charged bunches pass through each other...

Both E and B fields are *attractive* to the particles in the other bunch

If two bunches with the *same* sign pass through each other...

Both E and B fields are *repulsive* to the particles in the other bunch

In either case, the forces add

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$$\vec{F} = -\hat{r} \frac{e^2}{2\pi\epsilon_0 r} \frac{N}{L} \left(1 - e^{-r^2/2\sigma^2}\right) (1 + \beta^2) \approx 2$$

$$\approx -\hat{r} \frac{e^2}{\pi\epsilon_0 r} \frac{N}{L} \left(1 - e^{-r^2/2\sigma^2}\right)$$

### Effective Length

Front of first bunch encounters front of second bunch

Front of first bunch exits second bunch.

$\frac{L}{2}$  "Effective length"

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$$\Delta x' = \frac{F_x}{vp} \Delta s = \frac{F_x}{vp} \left( \frac{L}{2} \right)$$

$$= -\frac{N_b e^2}{2\pi\epsilon_0 r \gamma \beta^2 m c^2 r} x \left( 1 - e^{-r^2/2\sigma^2} \right)$$

$\beta = 1$

$$\approx -\frac{2N_b r_0}{\gamma} \frac{x}{r^2} \left( 1 - e^{-r^2/2\sigma^2} \right)$$

$$= -\frac{2N_b r_0}{\gamma} \frac{x}{(x^2 + y^2)} \left( 1 - e^{-\frac{(x^2 + y^2)}{2\sigma^2}} \right)$$

Small x and y

$$\approx -\frac{N_b r_0}{\gamma \sigma^2} x = -\frac{1}{f_{eff}} x$$

$$\Delta y' \approx -\frac{N_b r_0}{\gamma \sigma^2} y = -\frac{1}{f_{eff}} y$$

$$\Delta v = \frac{1}{4\pi} \frac{\beta^*}{f_{eff}}$$

$$= \frac{N_b r_0}{4\pi} \frac{\beta^*}{\gamma \sigma^2} = \frac{N_b r_0}{4\pi \epsilon_N}$$

normalized emittance (protons)

$$\equiv \xi$$

"Tuneshift Parameter"

Maximum tuneshift for particles near center of bunch

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## Luminosity and Tuneshift

The total tuneshift will ultimately limit the performance of any collider, by driving the beam onto an unstable resonance. Values of on the order ~.02 are typically the limit. However, we have seen the somewhat surprising result that the tuneshift

$$\xi = \frac{N_b r_0}{2\pi\epsilon\gamma}$$

does not depend on  $\beta^*$ , but only on

$$\frac{N_b}{\epsilon} \equiv \text{"brightness"}$$

For a collider, we have

$$\mathcal{L} = \frac{fn_b N_b^2}{4\pi\sigma^2} = \frac{fn_b N_b^2}{4\pi \left( \frac{\beta^* \epsilon_N}{\gamma} \right)} = \frac{fn_b N_b \gamma}{r_0 \beta^*} \left( \frac{r_0}{4\pi} \frac{N_b}{\epsilon_N} \right)$$

$$= f \frac{n_b N_b \gamma}{r_0 \beta^*} \xi$$

We assume we will run the collider at the "tuneshift limit", in which case we can increase luminosity by

- Making  $\beta^*$  as small as possible
- Increasing  $N_b$  and  $\epsilon$  proportionally.