



TRICKS OF THE TRADE

Eric Prebys, UC Davis



USPAS Fundamentals, June 4-15, 2018

E. Prebys, Accelerator Fundamentals: Tricks of the Trade

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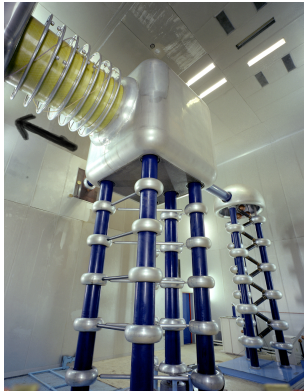
Multi-stage Acceleration

- Early synchrotrons had low energy injection and provided all the acceleration in a single stage.
- The energy range of a single synchrotron is limited by
 - An aperture large enough for the injected beam is unreasonably large at high field.
 - Hysteresis effects result in excessive nonlinear terms at low energy (very important for colliders)
- Typical range 10-20 for colliders, larger for fixed target
 - Fermilab Main Ring: 8-400 GeV (50x)
 - Fermilab Tevatron: 150-980 GeV (6.5x)
 - LHC: 400-7000 GeV (17x)
- The highest energy beams require multiple stages of acceleration, with high reliability at each stage
- How is this done?



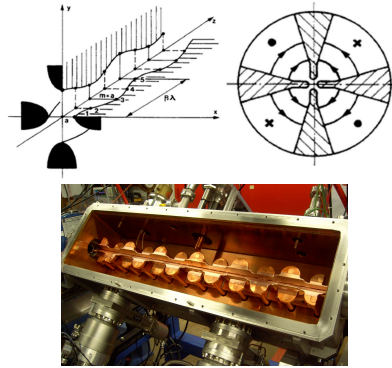
Initial Acceleration

Old: Static



Static acceleration from
Cockcroft-Walton.
FNAL = 750 keV
max ~1 MeV

New: RF Quadrupole (RFQ)

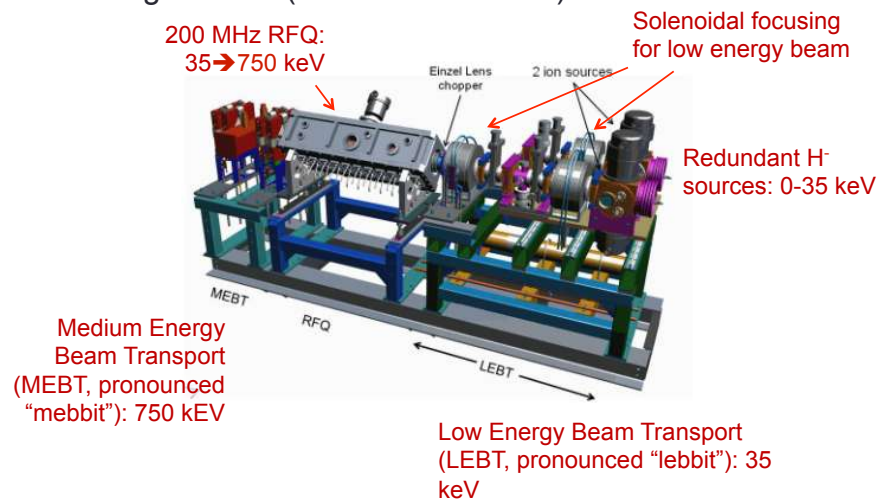


RF structure combines an electric
focusing quadrupole with a
longitudinal accelerating gradient.



Early Stages

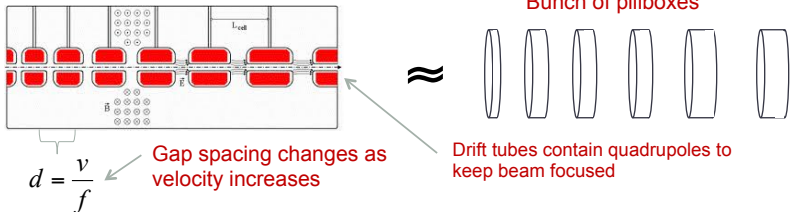
- The front end of any modern hadron accelerator looks something like this (Fermilab front end)



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Drift Tube (Alvarez) Cavity

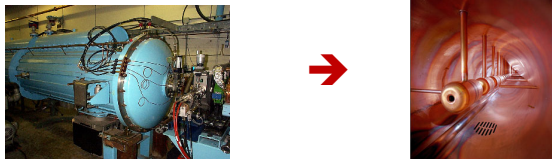
- Because the velocity is changing quickly, the first linac is generally a Drift Tube Linac (DTL), which can be beta-matched to the accelerating beam.
- Put conducting tubes in a larger pillbox, such that inside the tubes $E=0$



$d = \frac{v}{f}$ ← Gap spacing changes as velocity increases

Bunch of pillboxes

Drift tubes contain quadrupoles to keep beam focused




Fermilab low energy linac → Inside

- As energy gets higher, switch to “pi-cavities”, which are more efficient

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Linac -> Synchrotron Injection

- Eventually, the linear accelerator must inject into a synchrotron



- In order to maximize the intensity in the synchrotron, we can
 - Increase the linac current as high as possible and inject over one revolution
 - There are limits to linac current
 - Inject over multiple (N) revolutions of the synchrotron
 - Preferred method
- Unfortunately, Liouville's Theorem says we can't inject one beam on top of another
 - Electrons can be injected off orbit and will “cool” down to the equilibrium orbit via synchrotron radiation.
 - Protons can be injected a small, changing angle to “paint” phase space, resulting in increased emittance

$$\epsilon_s \geq N \epsilon_{LINAC}$$

Synchrotron emittance Linac emittance

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Ion (or Charge Exchange) Injection

Magnetic chicane pulsed to move beam out during injection

Circulating Beam

H^- beam from LINAC

Beam at injection

Stripping foil

- Instead of ionizing Hydrogen, and electron is added to create H^- , which is accelerated in the linac
- A pulsed chicane moves the circulating beam out during injection
- An injected H^- beam is bent in the opposite direction so it lies on top of the circulating beam
- The combined beam passes through a foil, which strips the two electrons, leaving a single, more intense proton beam.
- Fermilab was converted from proton to H^- during the 70's
- CERN *still* uses proton injection, but is in the process of upgrading (LINAC4 upgrade)

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Injection and Extraction

- We typically would like to extract (or inject) beam by switching a magnetic field on between two bunches (order ~ 10 -100 ns)

- Unfortunately, getting the required field in such a short time would result in prohibitively high inductive voltages, so we usually do it in two steps:

fast, weak "kicker"

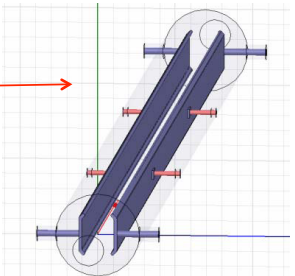
slower (or DC) extraction magnet with zero field on beam path.

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Extraction Hardware

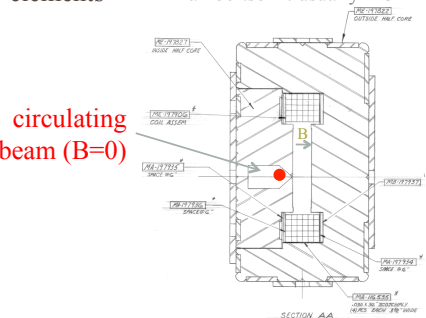
“Fast” kicker

- usually an impedance matched strip line, with or without ferrites

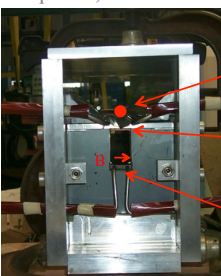


“Slow” extraction elements

“Lambertson”: usually DC



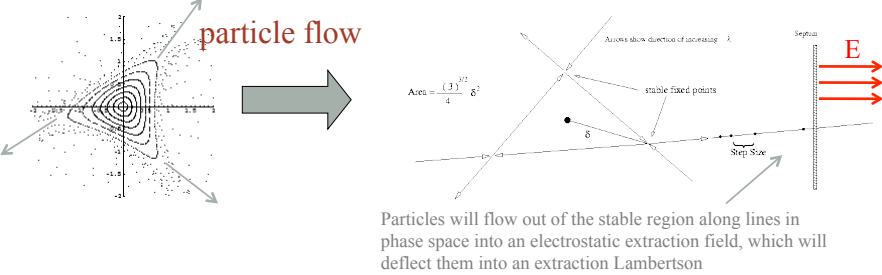
Septum: pulsed, but slower than the kicker



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Slow Extraction (not important for colliders)

- Sometimes fixed target experiments want beam delivered *slowly* (difficult)
- To do this, we generate a harmonic resonance
 - Usually sextupoles are used to create a 3rd order resonant instability



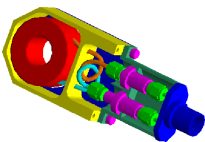
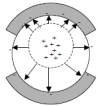
Particles will flow out of the stable region along lines in phase space into an electrostatic extraction field, which will deflect them into an extraction Lambertson


- Tune the instability so the escaping beam exactly fills the extraction gap between interceptions (3 times around for 3rd order)
 - Minimum inefficiency $\sim (\text{septum thickness})/(\text{gap size})$
 - Use electrostatic septum made of a plane of wires. Typical parameters
 - Septum thickness: .1 mm
 - Gap: 10 mm
 - Field: 80 kV

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Standard Beam Instrumentation

- Bunch/beam intensity are measured using inductive toroids
- Beam position is typically measured with beam position monitors (BPM's), which measure the induced signal on a opposing pickups
- Longitudinal profiles can be measured by introducing a resistor to measure the induced image current on the beam pipe -> Resistive Wall Monitor (RWM)

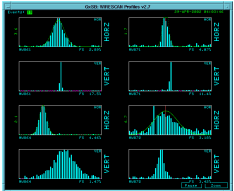



$$\Delta y \cong C \frac{I_{Top} - I_{Bottom}}{I_{Top} + I_{Bottom}}$$


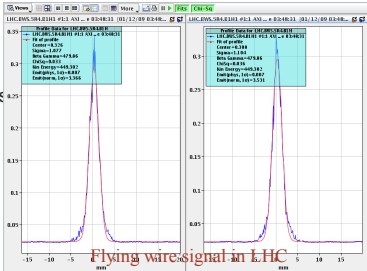
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Beam Instrumentation (cont'd)

- Beam profiles in beam lines can be measured using secondary emission multiwires (MW's)
- Can measure beam profiles in a circulating beam with a "flying wire scanner", which quickly passes a wire through and measures signal vs time to get profile
- Non-destructive measurements include
 - Ionization profile monitor (IPM): drift electrons or ions generated by beam passing through residual gas
 - Synchrotron light
 - Standard in electron machines
 - Also works in LHC



Beam profiles in MiniBooNE beam line

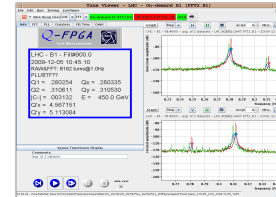


Flying wire signal in LHC

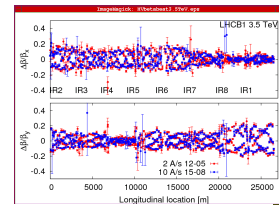


Measuring Lattice Parameters

- The fractional tune is measured by Fourier Transforming signals from the BPM's
 - Sometimes need to excite beam with a kicker

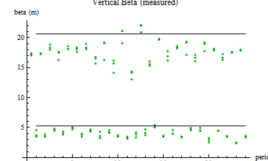


- Beta functions can be measured by exciting the beam and looking at distortions
 - Can use kicker or resonant ("AC") dipole



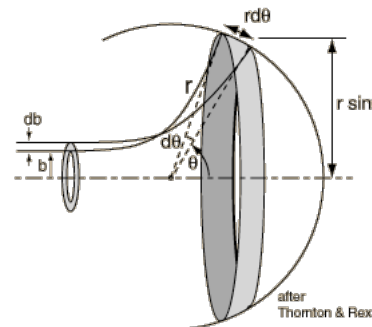
- Can also measure the by functions indirectly by varying a quad and measuring the tune shift

$$\Delta \nu = \frac{1}{4\pi} \frac{\partial \nu}{\partial k}$$



Understanding Cross Sections

- The formalism of cross sections was derived for Rutherford scattering, in which a scattered solid angle was mapped to an incident cross sectional area

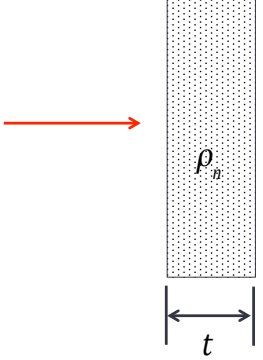


- It can be generalized to represent the probability of any sort of particle interaction

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Thin target approximation

- If the total probability that a particle will interact is small, then the probability of a particular interaction for one incident particle, will be given by



$$R = \sigma \times (\text{number of particles per cross-sectional area})$$

$$= \sigma \rho_n t$$

Target thickness [L]

Target number density [L⁻³]

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Luminosity

The relationship of the beam to the rate of observed physics processes is given by the "Luminosity"

Rate $\rightarrow R = L\sigma$

"Luminosity" \nearrow Cross-section ("physics") \nwarrow

Standard unit for Luminosity is cm⁻²s⁻¹
 Standard unit of cross section is "barn"=10⁻²⁴ cm²
 Integrated luminosity is usually in barn⁻¹, where

$$\text{b}^{-1} = (1 \text{ sec}) \times (10^{24} \text{ cm}^{-2} \text{ s}^{-1})$$

nb⁻¹ = 10⁹ b⁻¹, fb⁻¹=10¹⁵ b⁻¹, etc

For (thin) fixed target:

$$R = N\rho_n t\sigma \Rightarrow L = N\rho_n t$$

Incident rate \nearrow Target thickness \nwarrow

Target number density \nearrow

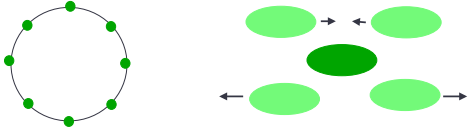
Example: MiniBooNe primary target:

$$L \approx 10^{37} \text{ cm}^{-2} \text{ s}^{-1}$$

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Colliding Beam Luminosity

Circulating beams typically “bunched” (number of interactions)



$$= \left(\frac{N_1}{A} \right) N_2 \sigma$$

Cross-sectional area of beam

Total Luminosity:

$$L = \left(\frac{N_1 N_2}{A} \right) r_b = \left(\frac{N_1 N_2}{A} \right) n \frac{c}{C}$$

crossing rate

Circumference of machine

Number of bunches

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Luminosity of Colliding Beams

- For equally intense Gaussian beams

Collision frequency

$$L = f \frac{N_b^2}{4\pi\sigma^2} R$$

Particles in a bunch

Geometrical factor:
- crossing angle
- hourglass effect

Transverse size (RMS)

prop. to energy

- Using $\sigma^2 = \frac{\beta^* \epsilon_N}{\gamma} \approx \frac{\beta^* \epsilon_N}{\gamma}$ we have

$$L = f_{rev} \frac{1}{4\pi} n_b N_b^2 \frac{\gamma}{\beta^* \epsilon_N} R$$

Revolution frequency

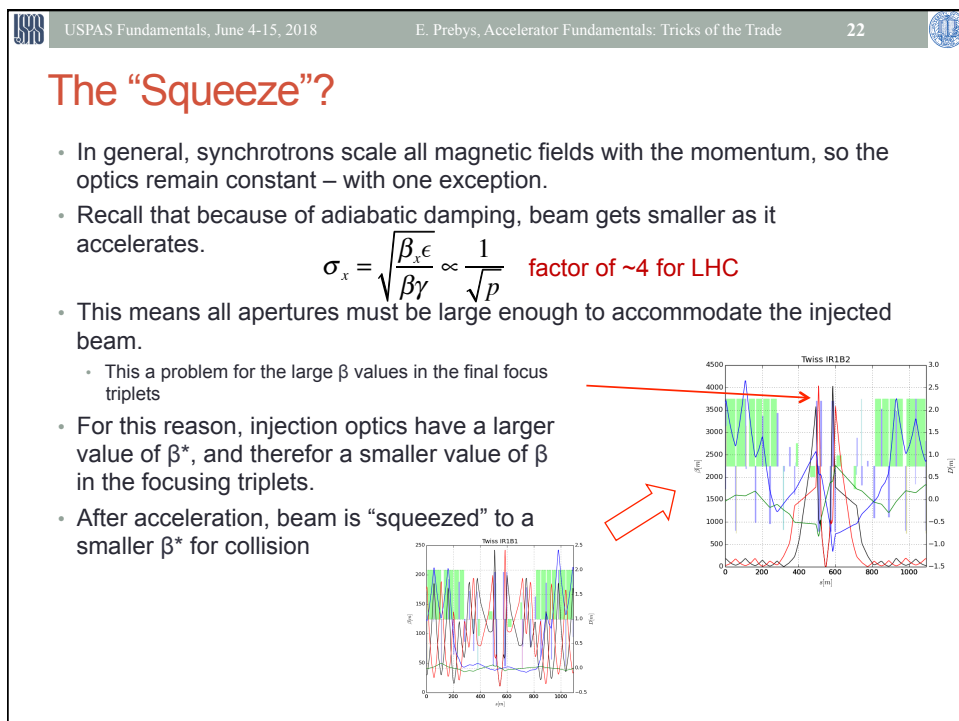
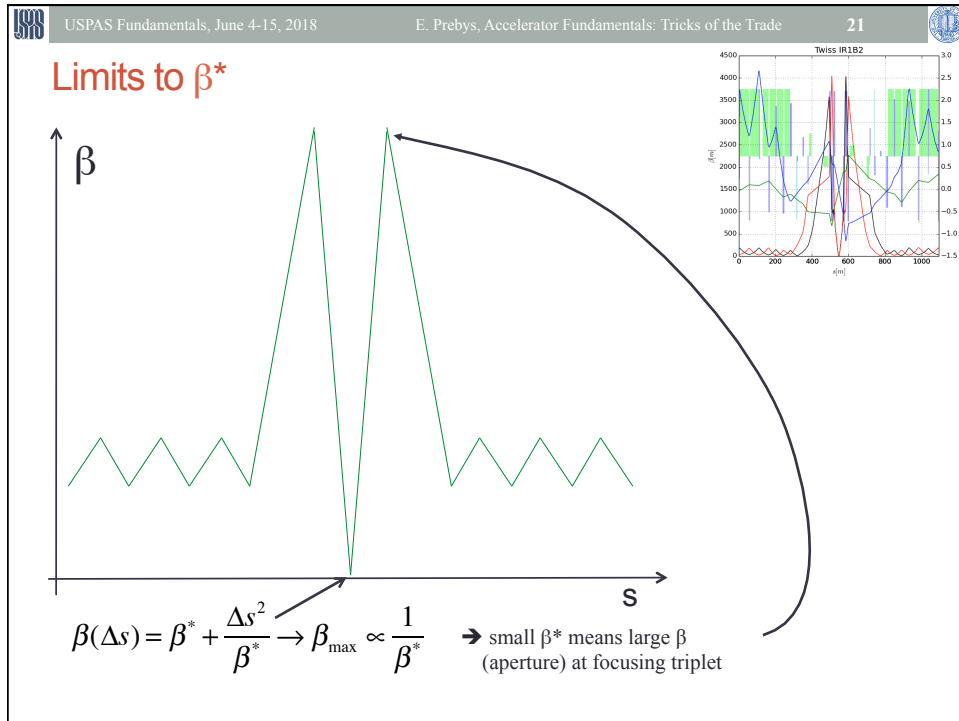
Number of bunches

Particles in bunch

Normalized emittance

Betatron function at collision point →
want a small β^* !


Record e+e- Luminosity (KEK-B):	$2.11 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$
Record p-pBar Luminosity (Tevatron):	$4.06 \times 10^{32} \text{ cm}^{-2}\text{s}^{-1}$
Record Hadronic Luminosity (LHC):	$2.06 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$



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Orbit correction

- Generally, beam lines or synchrotrons will have beam position monitors (BPM's) and correction dipoles (trims)



- We would like to use the trims to cancel out the effect of beamline imperfections, ie

$$-\Delta x_i = \sum A_{ij} \theta_j$$

Cancel displacement at BPM i due to imperfections

Setting of trim j

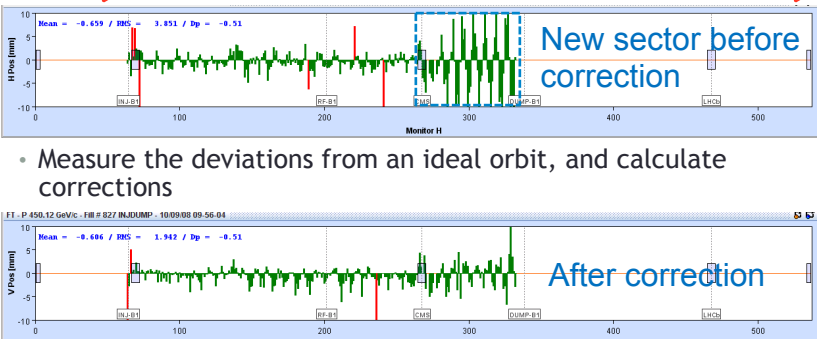
- Can express this as a matrix and invert to solve with standard techniques
 - If $n=m$, can just invert
 - If $n>m$, can minimize RMS

$$-\begin{pmatrix} \Delta x_0 \\ \Delta x_1 \\ \vdots \\ \Delta x_n \end{pmatrix} = \begin{pmatrix} A_{00} & A_{01} & \cdots & A_{0m} \\ A_{10} & A_{11} & \cdots & A_{1m} \\ \vdots & \vdots & \ddots & \vdots \\ A_{n0} & A_{n1} & & A_{nm} \end{pmatrix} \begin{pmatrix} \theta_0 \\ \theta_1 \\ \vdots \\ \theta_m \end{pmatrix}$$

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Example: First Beam through LHC (Sept. 10, 2008)

- General procedure
 - Proceed one octant at a time, closing collimators at the next point.



- Measure the deviations from an ideal orbit, and calculate corrections

- Might need to iterate a few times

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Beam Collimation and Machine Protection

- As beams get more intense, machine protection becomes very important
 - Full LHC energy ~ 150 sticks of dynamite!
- Beam halo is generally cleaned up through multi-stage collimation

The diagram illustrates the multi-stage collimation process for a proton beam. It starts with a beam having a 'Core' and 'Primary halo (p)'. The beam propagates through a 'Primary collimator' (CFC) with an impact parameter $\leq 1 \mu\text{m}$. This stage causes 'Unavoidable losses' and produces a 'Shower' of particles (e, p). The remaining beam has a 'Secondary halo' and passes through a 'Secondary collimator' (CFC), which also produces a 'Shower' (e, p). The beam then passes through an 'Absorber' (W/Cu), which produces a 'Tertiary halo'. Finally, it passes through another 'Absorber' (W/Cu) and 'Super-conducting magnets' before reaching 'SC magnets and particle physics exp.'. A text box notes: 'Without beam cleaning (collimators): Quasi immediate quench of super-conducting magnets (for higher intensities) and stop of physics. Required very good cleaning efficiency'.

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Secondary Beams

- When a proton beam strikes a target, the energy of the beam goes into particle production. Charged particles include (in \sim descending order of population)
 - π^\pm : Most of the energy
 - K^\pm : Charged particles containing a Strange quark
 - p: ordinary protons
 - e^\pm : These mostly come from neutral pions that immediately decay to two photons.
 - Antiprotons:
 - Other strange "hyperons"
- When an electron beam strikes a target, it makes mostly photons and e^\pm
 - Positron production targets can be very efficient.
- Generally, we design secondary beam lines to maximize acceptance of the species of beam we're looking for.



Special Case: Neutrino Beams

- Electron neutrinos are generally produced in nuclear reactors. High energy particle beams are used to produce primarily muon neutrinos in the reactions

$$\pi^- \rightarrow \mu^- \bar{\nu}_\mu \rightarrow (\nu_\mu e^- \bar{\nu}_e) \bar{\nu}_\mu$$

$$\pi^+ \rightarrow \mu^+ \nu_\mu \rightarrow (\bar{\nu}_\mu e^+ \nu_e) \nu_\mu$$

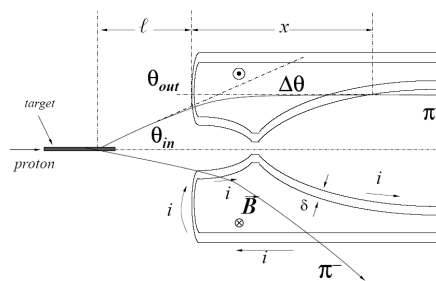
Leading particles

- Select correct neutrino species by focusing correct pion species



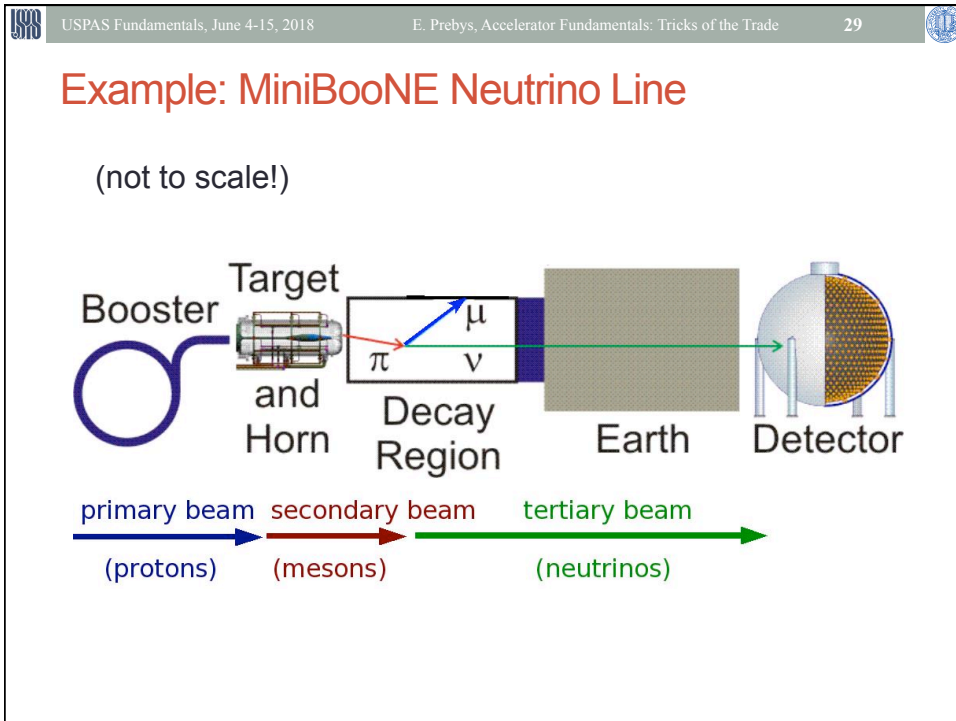
Neutrino Horns

- Neutrino horns work by producing a coaxial current so the correct sign pions are focused in *both* planes.



- This is the to decay

allow the pions



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Practical Considerations

- Neutrino horns operate in fierce radiation environments, and are pulsed with currents of several hundred kA.
- They require water cooling and sophisticated mechanical analyses.

Neutrino Horn Assembly at J-Parc

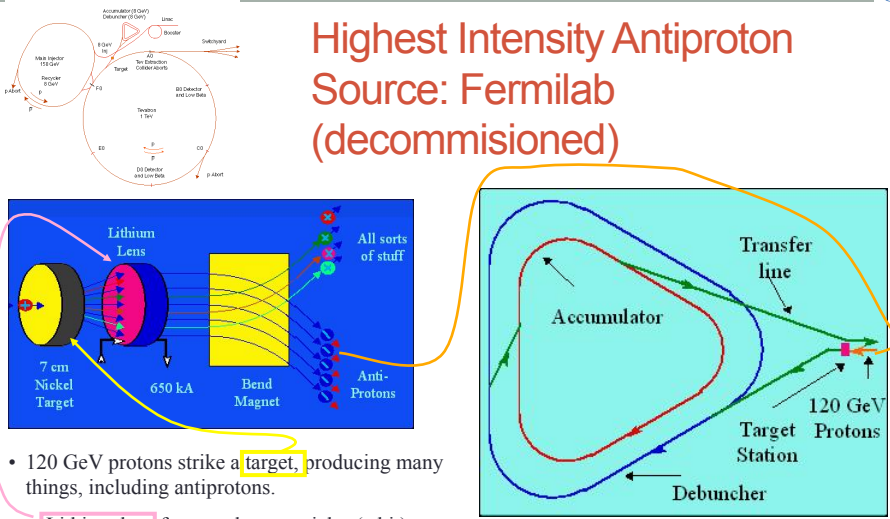
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Antiproton Beams

- Antiprotons are produced in very small numbers in proton collisions.
- In order to be useful, these must be captured and “cooled” (i.e. have their area in phase space reduced).
- Although high energy proton-antiproton colliders are a thing of the past (homework problem), anti-protons are still of great interest at low energy:
 - CERN LEAR facility
 - FAIR Facility in Germany.

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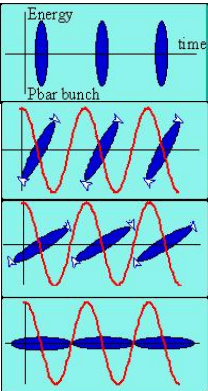
Highest Intensity Antiproton Source: Fermilab (decommissioned)



- 120 GeV protons strike a target, producing many things, including antiprotons.
- a Lithium lens focuses these particles (a bit)
- a bend magnet selects the negative particles around 8 GeV. Everything but antiprotons decays away.
- The antiproton ring consists of 2 parts
 - the Debuncher
 - the Accumulator.

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Antiproton Source – debunching



Particles enter with a **narrow** time spread and **broad** energy spread.

High (**low**) energy pbars take **more** (**less**) to go around...

...and the RF is phased so they are **decelerated** (**accelerated**),

resulting in a **narrow** energy spread and **broad** time spread.

At this point, the pBars are transferred to the accumulator, where they are “stacked”

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Stochastic cooling of antiprotons

- Positrons will naturally “cool” (approach a small equilibrium emittance) via synchrotron radiation.
- Antiprotons must rely on active cooling to be useful in colliders.
- Principle: consider a single particle which is off orbit. We can detect its deviation at one point, and correct it at another:
- But wait! If we apply this technique to an ensemble of particles, won't it just act on the centroid of the distribution? Yes, but...
- Stochastic cooling relies on “mixing”, the fact that particles of different momenta will slip in time and the sampled combinations will change.
- *Statistically*, the mean displacement will be dominated by the high amplitude particles and over time the distribution will cool.
- Simon Van der Meer won the Nobel Prize for this.

